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George Bendo

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| $\Box \oplus \leftrightarrow \sim \forall$ | 2011.0.00131.5 |) R Scl | 01:26:58.079 | -32:32:36.424 7 | 0.91 <mark>1</mark> 5 | 330.246.346.109 GHz | 2012-12-06 | 5 | 1.043 | 0.846 | 12m mosaic | 11.517 | 62.007 | Stars and stellar evolution | Asymptotic Giant Branch |
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- The sky viewer
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- Observation
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| $\bigcirc \Leftrightarrow \sim \checkmark \boxed{2011.0.00268.S} \qquad \text{Metallicity of a Submillimeter Galaxy at z=5}$ | S | Nagao, Tohru | De Breuck, Carlos; Hats 2013-02- | -09 <u>3</u> <u>4</u> LES | SS J0332-2756 |
| □ \oplus \leftrightarrow \sim \downarrow_{4} 2011.0.00454.5 (Why) Is CenA a source of Ultra High Energy Cosmic Rays: Shock acceleration, jet and UHECR composition | S | Nagar, Neil | Smith, Rory; Finlez, Car 2013-02- | -14 1 8 Ba | nd 6 CenA - CO knot S1 |
| $\bigcirc \Leftrightarrow \leftrightarrow \sim \pm$ 2011.0.00851S The Origin of the Destroyed Minor Planet at G29-38: a Main Belt or Kuiper Belt Analog? | S | Farihi, Jay | Greaves, Jane; Bonsor, 2013-02- | 14 1 8 G2 | 9-38 Band 6 RA=23: Run x2, G29-38 Band 7 RA=23: Run x5 |
| $\bigcirc \ \leftrightarrow \ \sim \ \downarrow_{s}$ Z011.0.00294.5 More than LESS: The first fully-identified submillimetre survey | S | Smail, Ian | Rix, Hans-Walter; Chap 2013-02- | -15 20 140 Tar | gets1-16, Targets1-16 Targets17-32, Targets112-126, Target: |
| □ ↔ ↔ ∠ 2011.0.00510.5 Probing the Molecular Outflows of the Coldest Known Object in the Universe: The Boomerang Nebula | S | Sahai, Raghvendra | Nyman, Lars-Ake; Vlem 2013-03- | -13 Z 6 B3 | 1 SB of 1 - Boomerang Nebula CO 1-0, B6 1 SB of 1 Boome |
| □ ↔ ↔ ∠ 2011.0.00131S Piecing the shell together: ALMA and the detached shell around R Scl | s | Maercker, Matthias | Ramstedt, Sofia; Paladi 2013-03- | -29 5 14 R S | Scl B3 Spec 1: Run x2, R Scl B6: Run x3, R Scl B7: Run x4 |
| ⊕ ↔ ∧ ↓ 2011.0.00808.5 Probing the vertical structure of Saturn's storm with ALMA | S | Cavalie, Thibault | Moreno, Raphael; Fouc 2013-04- | -23 0 4 GR | OUP_1_SB: Run directly after GROUP_2_SB, GROUP_1_SB: Ru |
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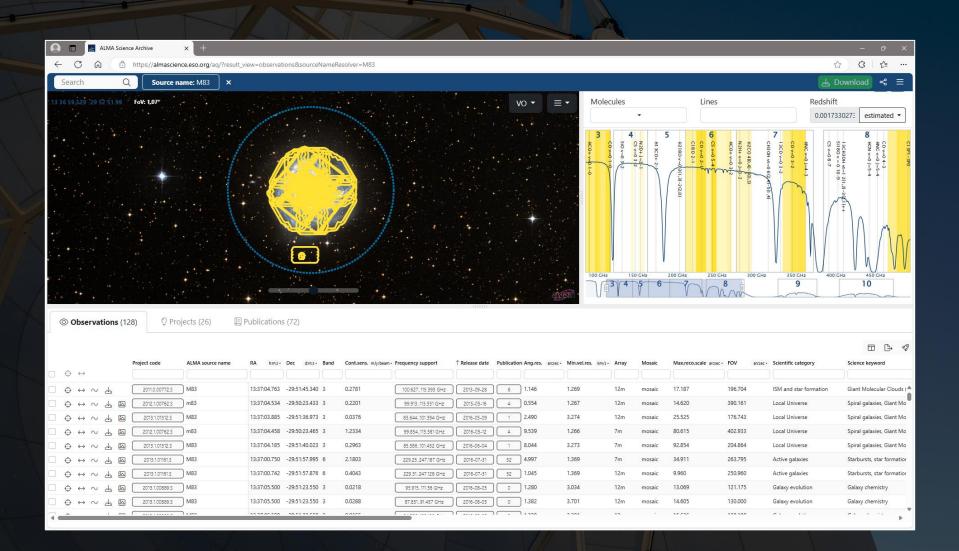
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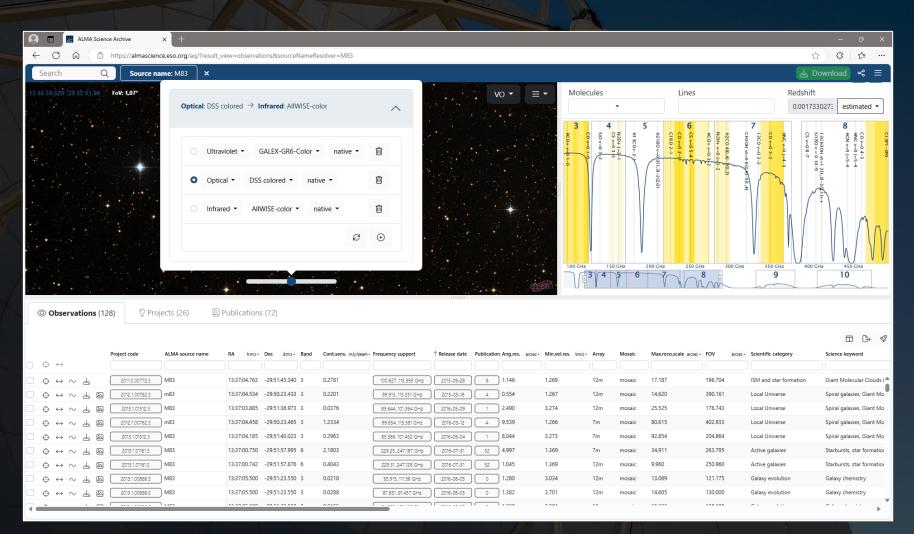
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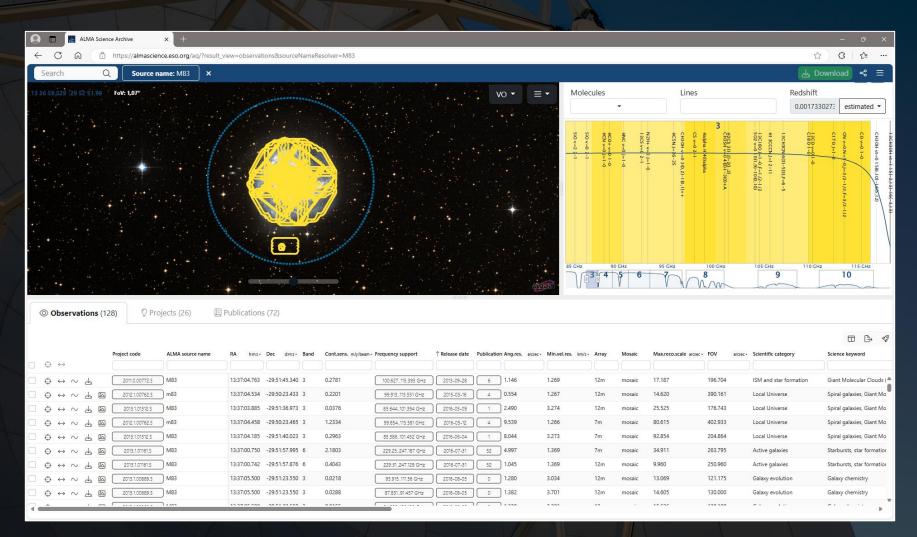
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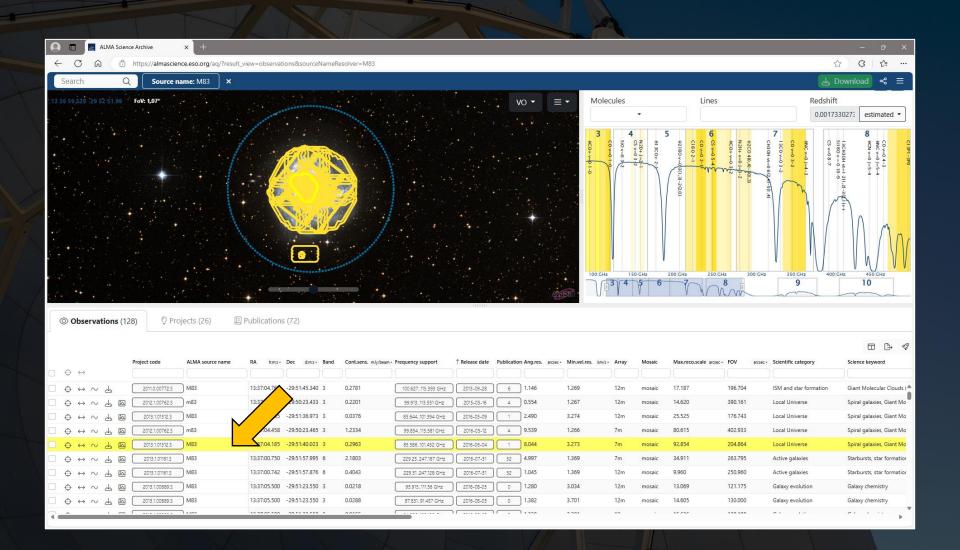
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Hovering over an entry in the results table will highlight the row, the field in the map panel, and the frequency ranges in the spectrum panel.



Hovering the cursor over items in boxes will reveal a pop-up window with extra information.

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| © Observations (124) | of M83 in 12C0 (1-0) enabled us to identify ~200 giant molecular clouds (GMCs) over wide range of galactic environments. Comparison with the HII regions indicated that progress of star formation strongly depends on the ratio between two time scales, manuellent argement between time. In addition, by adopting a star formation in us (F1-aku) which relates, SFR, galax mass, and the two time scales, excellent argement between the observation and the model were obtained for the radial distribution based analyses. Motivated by this finding, we aim to investigate the formation process of dense gas. In terms of it's environmental dependence, with the deep HCN observation. Since gas clouds denser than 1e4 crm-3 is known to be the basic unit of star formation, to verify the environmental dependence of SF-law, it is more essential to clarify the formation process of such dense gas. By comparing the deep HCN data with the CD data, and with the already available working hypothesis (time scale dependence of SF), we will address the formation process of dense gas. | r support | e date Publications | Ang.res. arcsec+ Min | .vel.res. km/z - Array | The second secon | B • C • Max.reco.scale |
| | Acknowledgement: This paper makes use of the following ALMA data: ADS/JAO.ALMA#2013.1.01312.S ALMA is a partnership of ESO | 7.115.393 GHz 2013-0 | 9-28 6 | 1.146 1.26 | 9 12m | mosaic | 17.187 |
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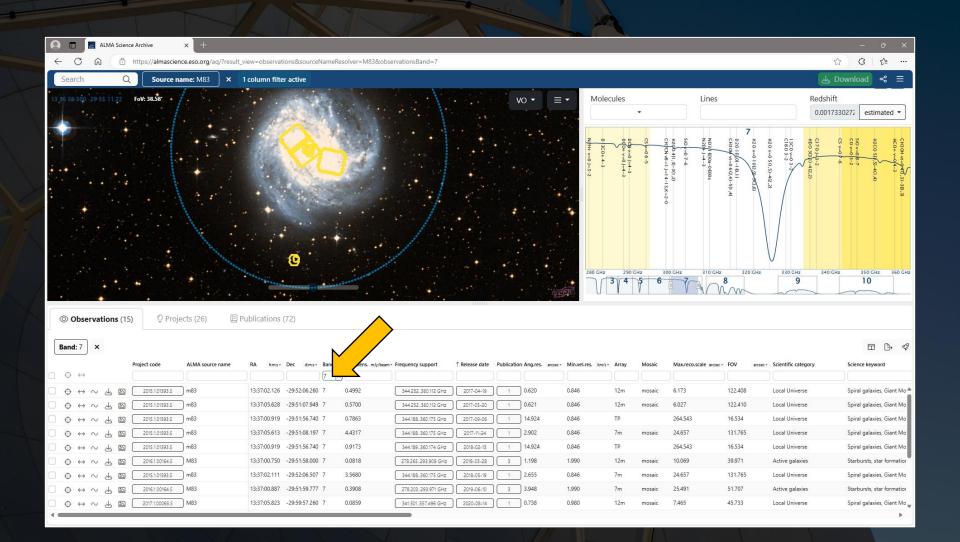
Clicking on the C symbol will launch CARTA, which can be used to inspect the data in more detail and even make measurements.

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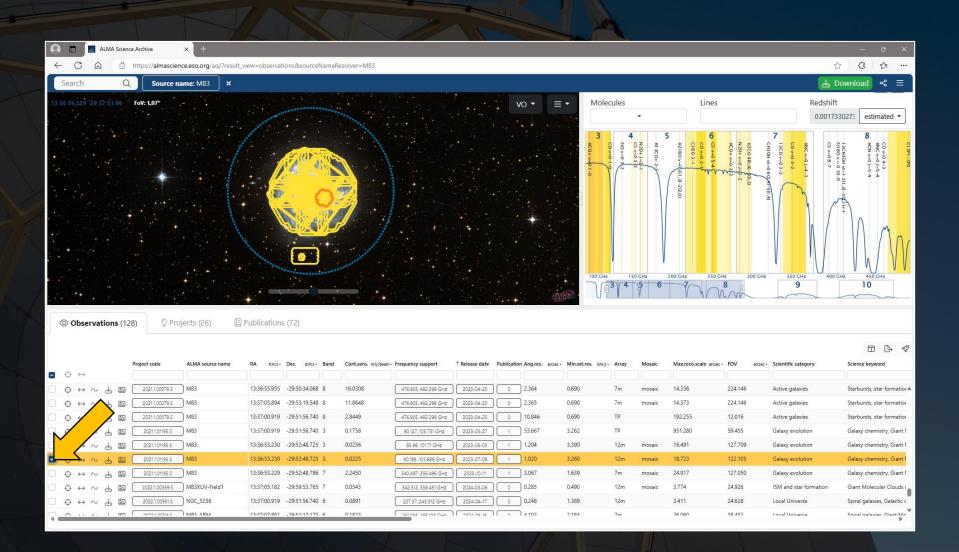
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| Φ | $\leftrightarrow \sim 2$ | y 🛛 (| 2015.1.00624.S |) M83 | 13:37:06.765 | -29:53:23.398 6 | 0.2940 | 213.927.231.155 GHz | 2018-06-28 | 0 | 0.135 0.367 | | 12m mosaic | 3.369 | 204.612 | Local Universe | Spiral galaxies, Galaxy ch |
| Φ | $\leftrightarrow \sim 2$ | L 🛛 (| 2013.1.01161.5 |) M83 | 13:37:00.750 | -29:51:58.000 6 | 0.1162 | 229.309.247.128 GHz | 2018-10-06 | 52 | 0.194 1.370 | 1 | 12m mosaic | 3.909 | 91.615 | Active galaxies | Starbursts, star formation |
| • | $\leftrightarrow \sim d$ | L (| 2023.1.01671.S |) M83 | 13:37:00.705 | -29:51:58.428 8 | 0.3125 | 478.006.493.741 GHz | 2025-09-13 | 0 | 0.226 2.368 | i. | 12m mosaic | 4.207 | 19.243 | Active galaxies | Starbursts, star formation |
| Φ | $\leftrightarrow \sim d$ | L 🖂 (| 2022.1.00951.S |) NGC_5236 | 13:37:00.919 | -29:51:56.740 6 | 0.0891 | 227.37.245.512 GHz | 2024-04-17 | 0 | 0.248 1.389 | 1 | 12m | 3.411 | 24.628 | Local Universe | Spiral galaxies, Galactic c |
| Φ | $\leftrightarrow \sim \epsilon$ | L 🖂 (| 2022.1.00359.5 |) M83XUV-Field1 | 13:37:05.182 | -29:59:53.765 7 | 0.0543 | 342.512.358.451 GHz | 2024-03-06 | 0 | 0.285 0.490 | | 12m mosaic | 3.774 | 24.926 | ISM and star formation | Giant Molecular Clouds (|
| Φ | $\leftrightarrow \sim d$ | L 🖂 (| 2015.1.01177.S |) m83 | 13:37:00.919 | -29:51:56.740 3 | 0.0115 | 85.604101.271 GHz | 2017-11-07 | 5 | 0.375 3.470 | ř. | 12m | 7.319 | 62.319 | Active galaxies | Starbursts, star formation |
| Φ | $\leftrightarrow \sim d$ | 5 🛛 (| 2013.1.01161.S |) M83 | 13:37:00.742 | -29:51:57.876 6 | 0.2194 | 229.309.247.128 GHz | 2016-10-07 | 52 | 0.496 1.370 | č. | 12m mosaic | 4.300 | 250.960 | Active galaxies | Starbursts, star formation |
| Φ | $\leftrightarrow \sim d$ | L 🖂 (| 2013.1.00861.S |) M83 | 13:37:03.967 | -29:59:47.584 6 | 0.3025 | 214.933234.1 GHz | 2016-11-19 | 3 | 0.552 2.515 | | 12m mosaic | 5.143 | 194.285 | ISM and star formation | Inter-Stellar Medium (ISI |
| • | | - (| |) ₂₂ | 12.27.04.524 | 20.00.00.400.0 | 0.0004 | | · · · · · · · · · · · · · · · · · · · | · · · · · · · · · · · · · · · · · · · | A.F.F. + 367 | | *** | 11000 | 200.474 | 1 111 ⁻ | · · · · · · · · · · · · · · · · · · · |

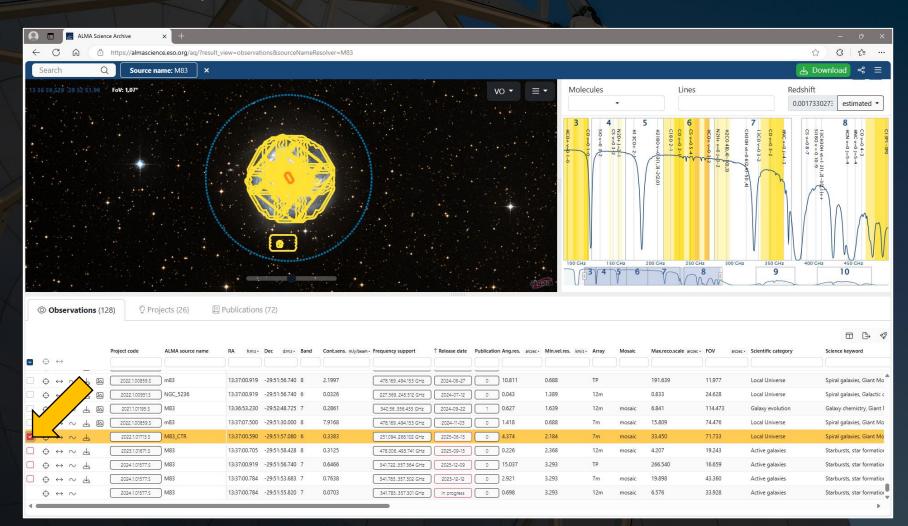
The results from a search can be sorted by any column. The results can also be further filtered.



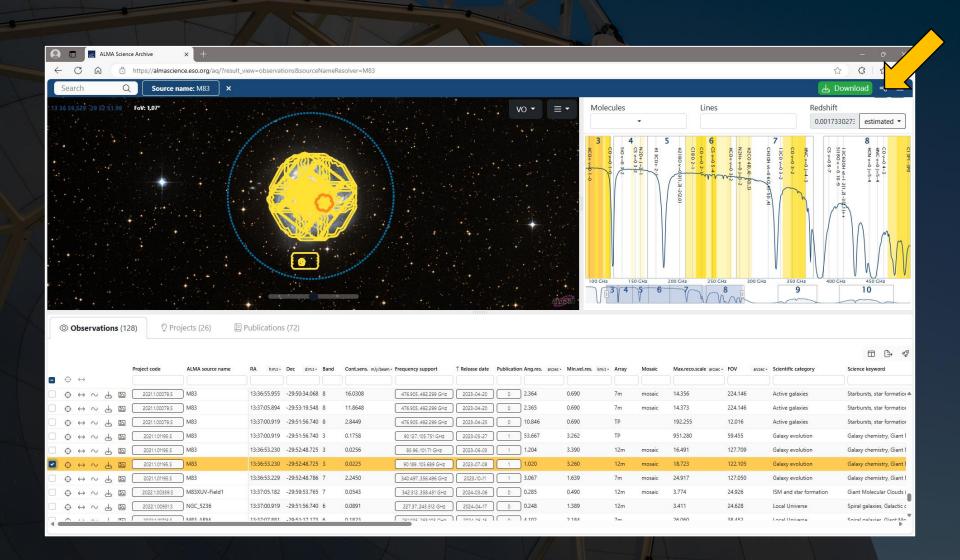
Clicking on the checkbox next to an observation will select the data for download. The row will change to orange as will the field in the map panel and the frequency range in the spectral plot.



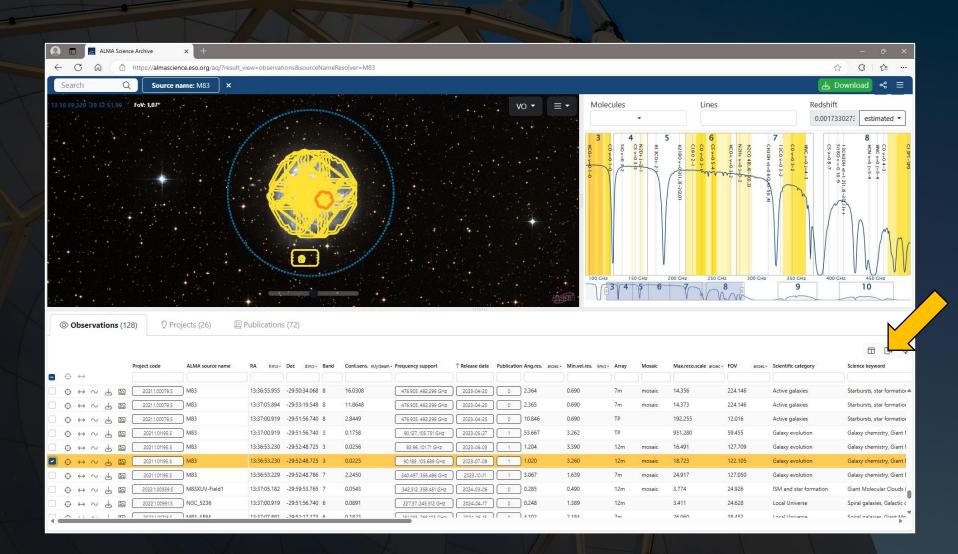
Proprietary data can be selected but cannot be downloaded. The checkbox will appear red when these data are selected. Other data (such as for programs where the observations are not yet complete or where the data are in QA3) cannot be selected.



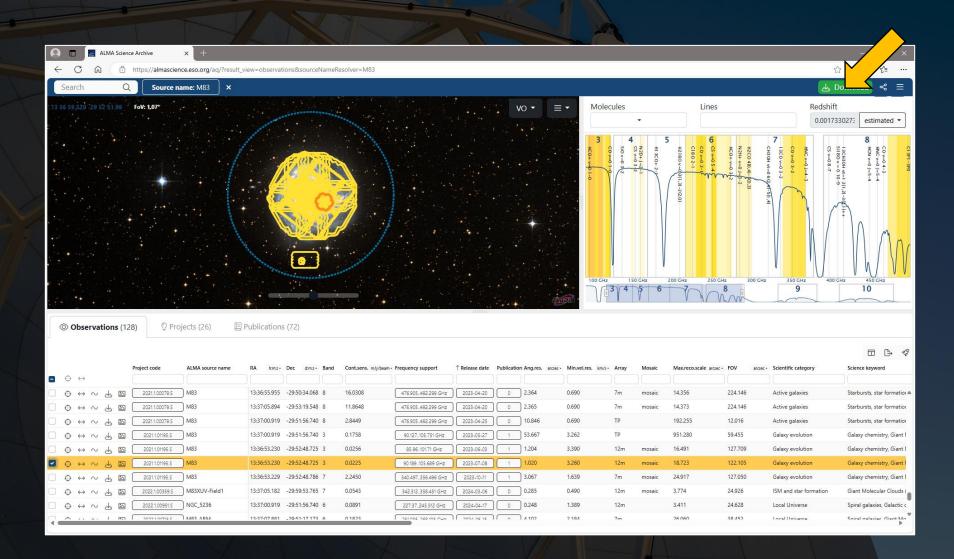
The interface has several other options as well. These include copying the link to the search results, selecting the columns that are displayed and saving the search results.



The interface has several other options as well. These include copying the link to the search results, selecting the columns that are displayed and saving the search results.



Selected data can be downloaded by clicking on the green "Explore and download" box at the top right. This will open a new display within the browser window listing the files associated with the selected dataset.



The Request Handler will display all of the files associated with the selected Scheduling Blocks.

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Each Member OUS (or Scheduling Block) may have the following files available for download:

product Final images and image cubes

auxiliary A file containing logs, quality assurance information, scripts, and calibration data

raw Raw visibility data

external Enhanced data products (including enhanced images or visibility data) created after the data delivery

The legacy version of this page is accessible through a link at the top, but it works poorly for projects with multiple Scheduling Blocks.

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| | | ✓ <u>2021.</u> | 1.01195.S uid A002 Xf96bbc X104b4.asdm.sdr | n.tar | (raw) | 38 GB | 2021.1.01195.S | uid://A001/X1590/Xd68 | uid://A001/X1590/Xd69 | |
| search for file type | 11 | _ | 1.01195.S uid A002 Xf934b1 X16672.asdm.sdr | | (raw) | 36 GB | 2021.1.01195.S | uid://A001/X1590/Xd68 | uid://A001/X1590/Xd69 | |
| | _ | ✓ <u>2021.</u> | 1.01195.S uid A002 Xf89be2 Xc821.asdm.sdm | tar | (raw) | 32 GB | 2021.1.01195.S | uid://A001/X1590/Xd68 | uid://A001/X1590/Xd69 | |
| application/x-gzip | 48 | ✓ <u>2021.</u> | 1.01195.S uid A002 Xf89be2 Xc393.asdm.sdm | <u>tar</u> | (raw) | 32 GB | 2021.1.01195.S | uid://A001/X1590/Xd68 | uid://A001/X1590/Xd69 | |
| text/xml | | | | | | | | | | |
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| image/x-fits | 13 | | | | | | | | | | | |
| text/plain | 12 | | | | | | | | | | | |
| application/tar | 11 | | | | | | | | | | | |
| application/pdf | 10 | | | | | | | | | | | |
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The download script can be executed in a Linux/Mac console to download data. The file must be made executable using chmod before doing this. When the file is executed, the data will be downloaded to the current directory.

If the script is interrupted, it is possible to restart the downloads from where they were stopped by restarting the script.

After downloading all of the files, the script will ask whether to unpack the files and whether to unpack the files into the original directory structure. Always select the option to unpack the files into the original directory structure.